

# On a possible solution to the Hubble tension

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# Hubble constant

- Hubble law:

$$v = H_0 \cdot d$$

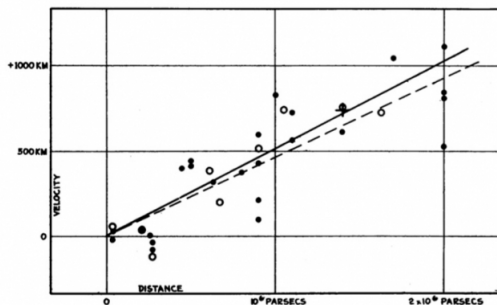


FIGURE 1  
Velocity-Distance Relation among Extra-Galactic Nebulae.

Edwin Hubble, Proceedings of the National Academy of Sciences, 1929

# Hubble constant

- Hubble law:

$$v = H_0 \cdot d$$

- cosmological Hubble law ( $\Omega_k = 0$ ):

$$(1+z) \cdot \int_0^z \frac{dz}{E(z)} = H_0 \cdot d_L,$$

where:

$$E(z) = (\Omega_m(1+z)^3 + \Omega_\Lambda)^{1/2}$$

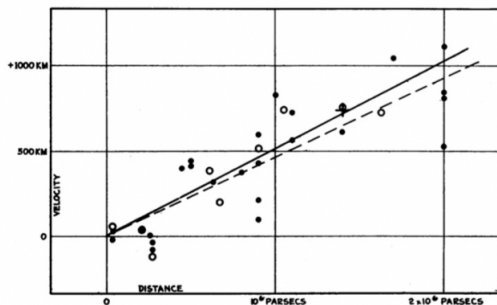
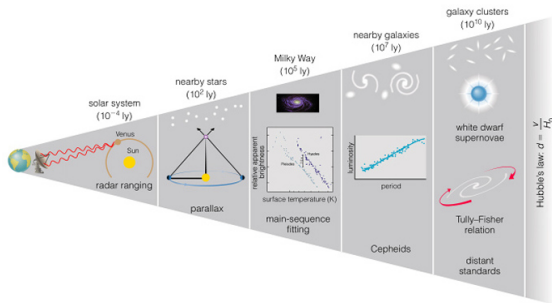


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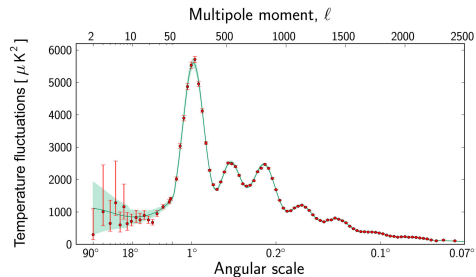
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# Hubble constant measurements

## Direct: distance ladder



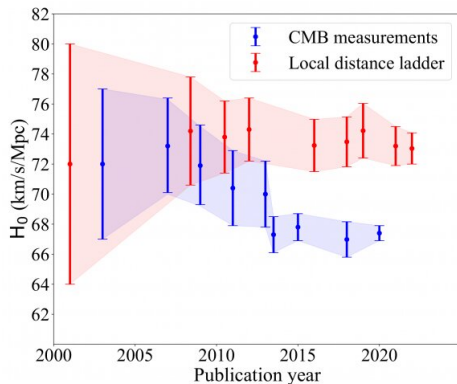
## Indirect: Cosmic Microwave Radiation



ESA and the Planck Collaboration

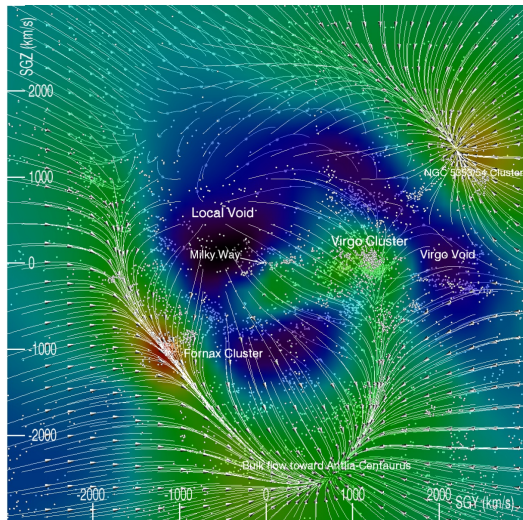
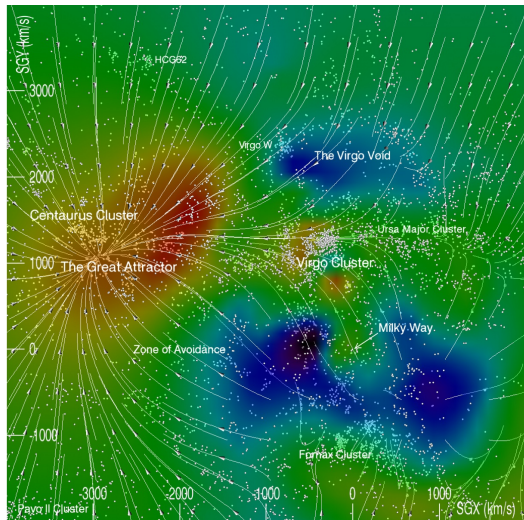
# Hubble tension

- the CMB-inferred Hubble constant is (Planck Collaboration 2020 *Astron. Astrophys*)  $H_0 \approx 67.4 \text{ km/s/Mpc}$
- local distance ladder measurements give (Riess et al 2022 *Astrophys. J. Lett.*)  $H_0 \approx 73 \text{ km/s/Mpc}$
- in the  $\Lambda$ CDM universe these values should be exactly the same



Perivolaropoulos, Skara, New Astronomy Review, 2022

# Local universe (Cosmicflows-4, Tully et al 2022, *ApJ*)



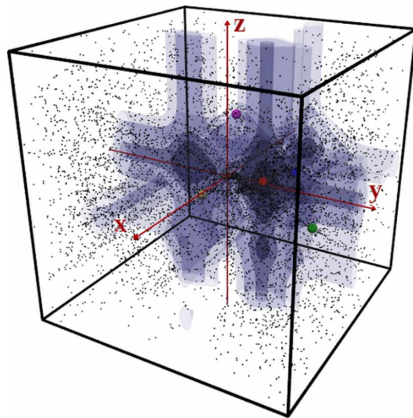
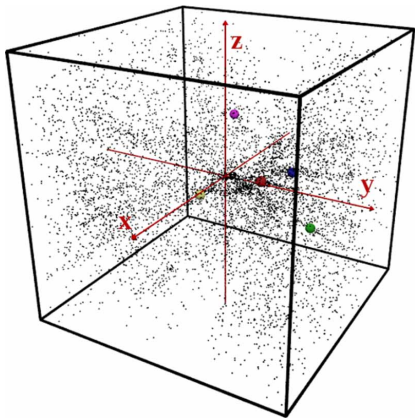
# Geometry of the local universe

Name	N	$X_i$	$Y_i$	$Z_i$
Virgo Cluster	164	-3.60	15.75	-0.65
Hydra Cluster	67	-32.62	27.95	-33.13
Fornax Cluster	51	-1.86	-14.49	-13.09
Centaurus Cluster	49	-36.11	15.83	-8.08
Pavo Cluster	18	-49.79	-23.49	10.83

- 2nd order perturbed metric of the local universe

$$ds^2 = -dt^2 + a(t)^2 \sum_{l=0}^2 \sum_{m=0}^l c_{ij}^{(l-m,m)}(x^\mu) \lambda^{l-m} k^m dx^i dx^j .$$

# Local density field





# Hubble constant: inhomogeneous model

**Standard  $\Lambda$ CDM approach**

**Inhomogeneous model**

# Hubble constant: inhomogeneous model

## Standard $\Lambda$ CDM approach

- collecting data e.g. type Ia supernovae light curves

## Inhomogeneous model

- mapping the local geometry and creating the mock catalog

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## Standard $\Lambda$ CDM approach

- collecting data e.g. type Ia supernovae light curves
- converting data to the  $d_L - z$  relation
- fitting the low-redshift formula

$$d_L = \frac{cz}{H_0} (1 + A z + B z^2 + O(z^3))$$

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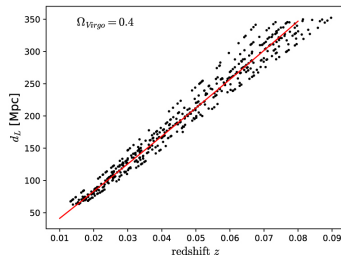
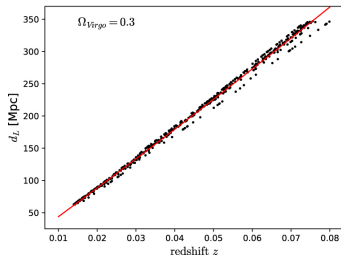
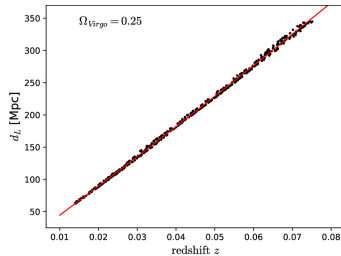
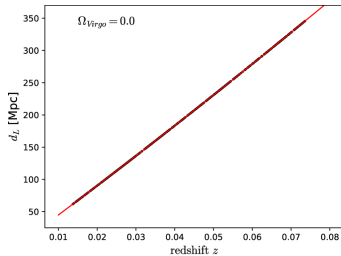
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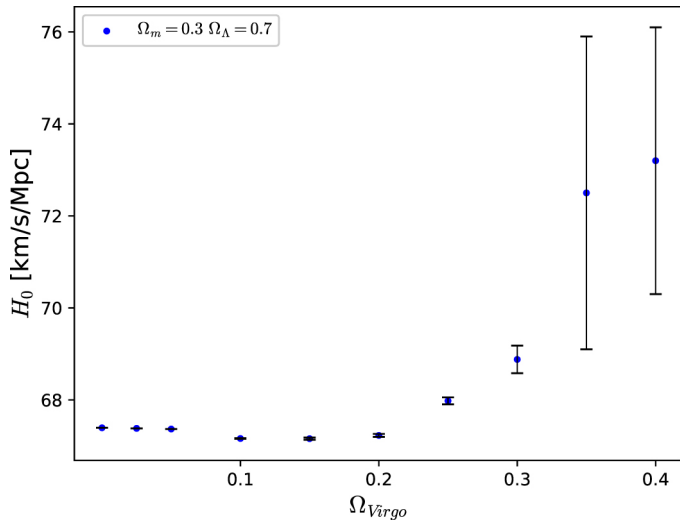
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# Mock data



# Hubble parameter in the 2nd order CPT



# Summary

- 2nd order CPT allows to probe higher than 1st order density contrast and its influence on the light propagation



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- results of simulations in the 2nd order CPT suggest that local density inhomogeneities increase the value of the Hubble constant if interpreted in the strictly  $\Lambda$  CDM context (consistent with Riess et al 2022 *Astrophys. J. Lett.*):

$\Omega_{Vir}$	<b>0.0</b>	0.1	0.15	0.2	0.25	0.3	0.35	<b>0.4</b>
$H_0$	<b>67.39</b>	67.1610	67.16	67.23	67.980	68.88	72.5	<b>73.2</b>
$\Delta H_0$	<b>1.7e-06</b>	0.00015	0.025	0.032	0.076	0.30	3.4	<b>2.9</b>

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- more details can be found in: Sikora, Ostrowski 2024, *Classical and Quantum Gravity*